

Two-level system with a large velocity gradient

David Neufeld

Corrected version: 2004 June 24th

Corrections highlighted in **yellow**

To test the ability of radiative transfer codes to model clouds with large velocity gradients, I suggest considering a two level case similar to that considered for the static case. Now, however, we introduce a large velocity gradient and – in the limit of large optical depth – can obtain an exact analytic solution.

As for the static case, we adopt an **excitation model** with

- Two-level ortho-water molecule (1_{01} and 1_{10} states)
- Collisional excitation of 557 GHz transition by ortho and para- H_2 in 3:1 ratio*
 - ➔ de-excitation rate coefficient, $q_{21} = 2.18 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - excitation rate coefficient, $q_{12} = 1.12 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - critical density, $n_{\text{cr}} = 1.59 \times 10^7 \text{ cm}^{-3}$

*Realistically, the ratio is almost certainly smaller than 3

and a **cloud model** with:

Outer radius, $R_{\text{out}} = 0.1 \text{ pc}$
 Inner radius = 0.001 pc
 H_2 density = 10^4 cm^{-3}
 Temperature = 40 K
 No external radiation

Now, however, we introduce a large **radial outflow velocity**

$$v_R = \alpha R, \text{ with } \alpha = 100 \text{ km/s/pc}$$

and adopt a Gaussian line width $\Delta v \ll v_R(R_{\text{out}}) = 10 \text{ km/s}$

The Sobolev optical depth is given by

$$\begin{aligned} \tau_S &= (f_1 - f_2) A_{21} \lambda^2 n(\text{o-}H_2O) / [8 \pi \alpha] \\ &= (1 - 2f_2) t \end{aligned} \tag{1}$$

where f_1 and $f_2 = 1 - f_1$ are the populations in the lower (1_{01}) and upper (1_{10}) states (as a fraction of the total ortho- H_2O abundance),

and where $t = x(\text{o-}H_2O) / 8.132 \times 10^{-9}$ and $x(\text{o-}H_2O)$ is the abundance of ortho-water
 The quantity t represents the Sobolev optical depth that would have resulted had all the water molecules been in the lower state.

Level populations

Except within a thin shell close to the surface (with a fractional thickness $\Delta v / v_R[R_{out}]$), the escape probability approximation is an excellent one, and the level populations are *independent of location* and given by the equation of statistical equilibrium

$$f_2/f_1 = q_{12} / (q_{21} + A_{21}\beta) = (q_{12}/q_{21}) / (1 + s\beta) \quad (2)$$

$$\text{where } \beta = (1 - \exp[-\tau_S]) / \tau_S$$

$$\text{and } s = n_{cr} / n[H_2] = 1588$$

In the limit $\tau_S \ll 1$, we obtain $\beta = 1$

$$\rightarrow f_2 = (q_{12}/q_{21}) / (1 + s) = 3.23 \times 10^{-4}$$

(as in the static, optically-thin case)

In the limit $\tau_S > \text{few}$, we may write $\beta \sim 1/\tau_S$

$$\rightarrow f_2 / (1 - f_2) = f_2/f_1 = (q_{12}/q_{21}) / (1 + s/\tau_S) = \tau_S u / (\tau_S + s) \quad (3)$$

$$\text{where } u = (q_{12}/q_{21}) = 0.512$$

Solving this equation simultaneously with equation (1), we obtain a quadratic equation for f_2 :

$$f_2 / (1 - f_2) = u t (1 - 2f_2) / [(1 - 2f_2) + s]$$

$$\rightarrow 2(u + 1) t f_2^2 - (3[u+1]t + s) f_2 + ut = 0$$

The solution is

$$f_2 = \{(3u + 1)t + s - [(1 - u)^2 t^2 + 2(3u + 1) st + s^2]^{1/2}\} / \{4(u+1)t\} \\ = \{2.536 t + 1588 - [0.238144 t^2 + 8054.336 t + 2521744]^{1/2}\} / \{6.048 t\} \quad (4)$$

which has the limiting behaviors

$$f_2 = u / (u + 1) = 0.339 \text{ (LTE value) in the limit } t \gg s \text{ (large abundance)}$$

and

$$f_2 = ut/s = 3.23 \times 10^{-4} t \text{ in the regime } \text{few} > t \gg s \text{ (intermediate abundance)}$$

NOTE: The above expressions only applies in the limit $t > \tau_s$, because we have used the approximation $\beta = (1 - \exp[-t/\tau_s]) / \tau_s \sim 1 / \tau_s$

However, for small t , and in the limit $s \gg 1$, we may approximate $t \sim \tau_s$ and write

$$f_2 = u/[s\beta] = u t / [s (1 - e^{-t})] = 3.23 \times 10^{-4} t / (1 - e^{-t}) \quad (5)$$

In Figure 1, the 1_{10} level population f_2 is plotted as a function of the ortho-water abundance, using equation (5) for abundances less than 3×10^{-8} (corresponding to $t < 3.7$), and equation (4) for greater abundances.

Line luminosity

The total line luminosity is

$$L = \mathcal{N}(\text{o-H}_2\text{O}) A_{21} h \nu f_2 \beta = 1.57 \times 10^{40} f_2 \beta x(\text{o-H}_2\text{O}) \text{ erg s}^{-1} \quad (6)$$

where $\mathcal{N}(\text{o-H}_2\text{O}) = 1.23 \times 10^{57} x(\text{o-H}_2\text{O})$ is the number of ortho-water molecules in the cloud.

The result is plotted in Figure 2.

In the limit $t \ll s$, we may write $f_2 \beta = u/s = 3.23 \times 10^{-4}$ and the luminosity is

$$L = 5.08 \times 10^{26} [x(\text{o-H}_2\text{O}) / 10^{-10}] \text{ erg s}^{-1} = 4.13 \times 10^{27} t \text{ erg s}^{-1}$$

(in agreement with the result previously obtained for the static case).

In the limit $t \gg s$, the gas is thermalized and optically-thick.

Under these conditions, $f_2 = 0.339$ and $\beta = 1 / [(1 - 2f_2) t] = 3.11 / t$ which implies $L = 1.35 \times 10^{31} \text{ erg s}^{-1}$. This is the maximum luminosity.



